

Mathematics of Complex and Nonlinear Phenomena



Towards a (2+1)D Generalised Hydrodynamics

ISIN2023 Northumbria University

Thibault Bonnemain, 28th April 2023 [Based on preliminary work with Benjamin Doyon]

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)
 - Bose-Einstein condensates (two-fluid model)

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)
 - Bose-Einstein condensates (two-fluid model)
 - Magneto hydrodynamics (plasma)

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)
 - Bose-Einstein condensates (two-fluid model)
 - Magneto hydrodynamics (plasma)
 - Relativistic hydrodynamics (superdense neutron star)

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)
 - Bose-Einstein condensates (two-fluid model)
 - Magneto hydrodynamics (plasma)
 - Relativistic hydrodynamics (superdense neutron star)

\Rightarrow Generalised hydrodynamics (integrable systems)

[Castro-Alvaredo, Doyon, Yoshimura (2016)] [Bertini, Collura, De Nardis, Fagotti (2016)]

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)

\Rightarrow Generalised hydrodynamics (integrable systems)

[Castro-Alvaredo, Doyon, Yoshimura (2016)] [Bertini, Collura, De Nardis, Fagotti (2016)]

- Derived from an underlying microscopic model:
 - \Rightarrow field theories or many-particle systems

- Hydrodynamics is everywhere in physics:
 - Fluid dynamics (simple fluids Euler 1757)

\Rightarrow Generalised hydrodynamics (integrable systems)

[Castro-Alvaredo, Doyon, Yoshimura (2016)] [Bertini, Collura, De Nardis, Fagotti (2016)]

- Derived from an underlying microscopic model:
 - \Rightarrow field theories or many-particle systems
- Main ingredients:

 \Rightarrow local conservation laws + propagation of local "equilibrium"

• N particles on a circle of perimeter L





- N particles on a circle of perimeter L
- Conservation laws



 (x_j, p_j)

- N particles on a circle of perimeter L
- Conservation laws

$$N \quad , \qquad P = \sum_{j=1}^{N} p_j \quad , \qquad E = \sum_{j=1}^{N} \frac{p_j^2}{2} + \sum_{i \neq j} V(x_i - x_j)$$

so that

• Local densities

$$q_0(x) = \sum_{j=1}^N \delta(x - x_j)$$

$$q_1(x) = \sum_{j=1}^N \delta(x - x_j) p_j$$

$$q_2(x) = \sum_{j=1}^N \delta(x - x_j) \left[\frac{p_j^2}{2} + \sum_{i \neq j} V(x_i - x_j) \right]$$

$$N = \int_0^L dx \ q_0(x)$$
$$P = \int_0^L dx \ q_1(x)$$
$$E = \int_0^L dx \ q_2(x)$$

 (x_j, p_j)

$$\mathbf{V} = \int_0^L \mathrm{d}x \ q_0(x)$$
$$\mathbf{P} = \int_0^L \mathrm{d}x \ q_1(x)$$

- N particles on a circle of perimeter L
- Conservation laws

$$N$$
 , $P = \sum_{j=1}^{N} p_j$, $E = \sum_{j=1}^{N} \frac{p_j^2}{2} + \sum_{i \neq j} V(x_i - x_j)$

• Local densities

$$N = \int_0^L dx \ q_0(x) \quad , \qquad P = \int_0^L dx \ q_1(x) \quad , \qquad E = \int_0^L dx \ q_2(x)$$

• Local conservation laws

Currents

 $\mathrm{d}x$

 (x_j, p_j)

$$\partial_t q_m(x,t) + \partial_x j_m(x,t) = 0$$
, $m = 0, 1, 2$.

- N particles on a circle of perimeter L
- Conservation laws

$$N$$
 , $P = \sum_{j=1}^{N} p_j$, $E = \sum_{j=1}^{N} \frac{p_j^2}{2} + \sum_{i \neq j} V(x_i - x_j)$

 (x_j, p_j)

Currents

 $\mathrm{d}x$

• Local densities

$$N = \int_0^L dx \ q_0(x) \quad , \qquad P = \int_0^L dx \ q_1(x) \quad , \qquad E = \int_0^L dx \ q_2(x)$$

• Local conservation laws

$$\partial_t q_m(x,t) + \partial_x j_m(x,t) = 0$$
, $m = 0, 1, 2$.

Functions on phase space or field operators

• Boltzmann 1868: micro-canonical ensemble in long time limit

4

• Boltzmann 1868: micro-canonical ensemble in long time limit

$$\triangleq \quad \text{Gibbs ensembles (GE):} \quad \rho \propto \exp[-\beta(E - \mu N - \nu P)]$$

4

• Boltzmann 1868: micro-canonical ensemble in long time limit

 $\triangleq \quad \text{Gibbs ensembles (GE):} \quad \rho \propto \exp[-\beta(E - \mu N - \nu P)]$

• Hydrodynamic principle: separation of scales and propagation of local GE



[Doyon: Lecture Notes (2020)]

• Boltzmann 1868: micro-canonical ensemble in long time limit

≙ Gibbs ensembles (GE): $\rho \propto \exp[-\beta(E - \mu N - \nu P)]$

• Hydrodynamic principle: separation of scales and propagation of local GE



[Doyon: Lecture Notes (2020)]

• Boltzmann 1868: micro-canonical ensemble in long time limit

 $\triangleq \quad \text{Gibbs ensembles (GE):} \quad \rho \propto \exp[-\beta(E - \mu N - \nu P)]$

• Hydrodynamic principle: separation of scales and propagation of local GE



• Boltzmann 1868: micro-canonical ensemble in long time limit

 $\triangleq \quad \text{Gibbs ensembles (GE):} \quad \rho \propto \exp[-\beta(E - \mu N - \nu P)]$

• Hydrodynamic principle: separation of scales and propagation of local GE



 \Rightarrow Local GE averages

 $\bar{q}_m(x,t)$ and $\bar{j}_m(x,t)$

 \cdot (approx) Meso conservation law

 $\partial_t \bar{q}_m(x,t) + \partial_x \bar{j}_m(x,t) = 0$

[Doyon: Lecture Notes (2020)]

• Boltzmann 1868: micro-canonical ensemble in long time limit

 $\triangleq \quad \text{Gibbs ensembles (GE):} \quad \rho \propto \exp[-\beta(E - \mu N - \nu P)]$

• Hydrodynamic principle: separation of scales and propagation of local GE



4

• Boltzmann 1868: micro-canonical ensemble in long time limit

$$\triangleq \quad \text{Generalised Gibbs ensembles (GE):} \quad \rho \propto e^{-\sum_{n=0}^{\infty} \beta_n Q_n}$$

 $\bullet\,$ Hydrodynamic principle: separation of scales and propagation of local GGE



[Doyon: Lecture Notes (2020)]

 \Rightarrow Local GGE averages

 $\bar{q}_m(x,t)$ and $\bar{j}_m(x,t)$

 $\Rightarrow \quad (\text{approx}) \text{ Meso conservation law} \\ \partial_t \bar{q}_m(x,t) + \partial_x \bar{j}_m(x,t) = 0 \\ \text{Functions of } \{\bar{q}_n\} \text{'s !} \end{aligned}$

4

• KdV: integrable nonlinear dispersive PDE

$$\partial_t u + 6u\partial_x u + \partial_x^3 u = 0$$

• KdV: integrable nonlinear dispersive PDE



Х

 $\partial_t u + 6u\partial_x u + \partial_x^3 u = 0 \; .$



Example of KdV

soliton gas

5

• KdV: integrable nonlinear dispersive PDE

$$\partial_t u + 6u\partial_x u + \partial_x^3 u = 0 \; .$$



• KdV: integrable nonlinear dispersive PDE

$$\partial_t u + 6u\partial_x u + \partial_x^3 u = 0 \; .$$

• Multisoliton solution

$$u_N \sim \sum_{i=1}^N 2\eta_i^2 \operatorname{sech}^2 \left[\eta_i \left(x - 4\eta_i^2 t - x_i^{\pm} \right) \right] \quad \text{as} \quad t \to \pm \infty.$$

N solitons



[Zakharov (1971)]

• KdV: integrable nonlinear dispersive PDE

$$\partial_t u + 6u\partial_x u + \partial_x^3 u = 0 \; .$$

• Multisoliton solution

$$u_N \sim \sum_{i=1}^N 2\eta_i^2 \operatorname{sech}^2 \left[\eta_i \left(x - 4\eta_i^2 t - x_i^{\pm} \right) \right] \quad \text{as} \quad t \to \pm \infty.$$
Action coordinate
Angle coordinate

N solitons



• KdV: integrable nonlinear dispersive PDE

$$\partial_t u + 6u\partial_x u + \partial_x^3 u = 0 \; .$$

• Multisoliton solution





Scattering is elastic and 2-body factorisable

N solitons



• KdV: integrable nonlinear dispersive PDE

$$\partial_t u + 6u\partial_x u + \partial_x^3 u = 0$$

• Multisoliton solution

$$u_N \sim \sum_{i=1}^N 2\eta_i^2 \operatorname{sech}^2 \left[\eta_i \left(x - 4\eta_i^2 t - x_i^{\pm} \right) \right] \quad \text{as} \quad t \to \pm \infty.$$
Action coordinate
Angle coordinate

• Relation between asymptotic states given by scattering shift

$$\sum_{i=1}^{t} \frac{\operatorname{sgn}(\eta_{i} - \eta_{j})}{\eta_{i}} \ln \left| \frac{\eta_{i} + \eta_{j}}{\eta_{i} - \eta_{j}} \right|$$
[Lax (1968)]

N solitons



5

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

Soliton bare velocity

$$p(\eta) = 4\eta^2$$

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

Soliton bare velocity

Generalised Gibbs weights

$$p(\eta) = 4\eta^2$$
 e.g.

$$w(\eta) = \sum_{k} \beta_k h_k(\eta)$$

 $h_n(\eta) = Q_n$ for a single soliton η

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

Soliton bare velocity

Generalised Gibbs weights

$$p(\eta) = 4\eta^2$$
 e.g

g.
$$w(\eta) = \sum_k \beta_k \eta^{2k+1}$$

 $h_n(\eta) = Q_n$ for a single soliton η

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

Soliton bare velocity

Generalised Gibbs weights

Constraint / Entropy

$$p(\eta) = 4\eta^2$$

e.g. $w(\eta) = \sum_k \beta_k \eta^{2k+1}$

 $h_n(\eta) = Q_n$ for a single soliton η

6

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

• Thermodynamic limit $L \to \infty$

$$\mathcal{Z} \asymp \exp\left(-L\mathcal{F}\right) , \qquad \mathcal{F} = -\int_{\Gamma} \frac{\eta \mathrm{d}\eta}{\sigma(\eta)}$$



• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

• Thermodynamic limit $L \to \infty$

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

• Thermodynamic limit $L \to \infty$

6

• NDR of soliton gases

$$\sigma(\eta)\rho(\eta) = \eta - \int_{\Gamma} d\mu \ \rho(\mu) \log \left| \frac{\eta + \mu}{\eta - \mu} \right|$$

 $\rho(\eta) d\eta dx = \#$ of solitons in $[x, x + dx] \times [\eta, \eta + d\eta]$ **Density of States**

[El (2003)]
Thermodynamics

• Partition function

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}p(\eta_i)}{2\pi} \mathrm{d}x_i^- \exp\left[-\sum_{i=1}^N w(\eta_i)\right] \chi\left(u_N(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

• Thermodynamic limit $L \to \infty$

• NDR of soliton gases

$$\sigma(\eta)\rho(\eta) = \eta - \int_{\Gamma} d\mu \ \rho(\mu) \log \left| \frac{\eta + \mu}{\eta - \mu} \right|$$

 $\rho(\eta) d\eta dx = \# \text{ of solitons in } [x, x + dx] \times [\eta, \eta + d\eta]$ Density of States
[EI (2003)]

• Alternative interpretation

NTDD

$$\frac{\mathrm{d}x^{-}(\eta)}{\mathrm{d}x} = \frac{\sigma(\eta)\rho(\eta)}{\eta}$$

Change of metric

• Integrability: infinite number of conservation laws

$$\partial_t q_n + \partial_x j_n = 0$$

• Integrability: infinite number of conservation laws

 $\partial_t q_n + \partial_x j_n = 0$

• Hydrodynamic approximation: separation of scales

 $\langle o(x,t) \rangle \approx \langle o \rangle_{\{\beta_n(x,t)\}} \equiv \bar{o}_n(x,t)$

• Integrability: infinite number of conservation laws

 $\partial_t q_n + \partial_x j_n = 0$

• Hydrodynamic approximation: separation of scales

(7)

• Integrability: infinite number of conservation laws

 $\partial_t q_n + \partial_x j_n = 0$

• Hydrodynamic approximation: separation of scales

(7)

• Integrability: infinite number of conservation laws

 $\partial_t q_n + \partial_x j_n = 0$

• Hydrodynamic approximation: separation of scales

• Integrability: infinite number of conservation laws

$$\partial_t q_n + \partial_x j_n = 0$$

• Hydrodynamic approximation: separation of scales

• Integrability: infinite number of conservation laws

 $\partial_t q_n + \partial_x j_n = 0$

• Hydrodynamic approximation: separation of scales

$$\langle o(x,t) \rangle \approx \langle o \rangle_{\{\beta_n(x,t)\}} \equiv \bar{o}_n(x,t)$$



Fluid cell average

$$\partial_t \rho(k; x, t) + \partial_x \left[v^{\text{eff}}(k; x, t) \rho(k; x, t) \right] = 0$$
$$v^{\text{eff}}(k) = 4\eta^2 + \frac{1}{\eta} \int_{\Gamma} \log \left| \frac{\eta + \mu}{\eta - \mu} \right| \rho(\mu) [v^{\text{eff}}(\eta) - v^{\text{eff}}(\mu)] d\mu$$

A type of (2+1)d GHD featuring line solitons

• Inspired by the phenomenology of the KP equation

$$\partial_x(\partial_t u + 6u\partial_x u + \partial_{xxx}u) \pm 3\partial_{yy}u = 0$$



A type of (2+1)d GHD featuring line solitons

• Inspired by the phenomenology of the KP equation

$$\partial_x(\partial_t u + 6u\partial_x u + \partial_{xxx}u) \pm 3\partial_{yy}u = 0$$



• Focus on gas of "lines" to make the jump from (1+1) to (2+1)D easier

A type of (2+1)d GHD featuring line solitons

• Inspired by the phenomenology of the KP equation

$$\partial_x(\partial_t u + 6u\partial_x u + \partial_{xxx}u) \pm 3\partial_{yy}u = 0$$



- Focus on gas of "lines" to make the jump from (1+1) to (2+1)D easier
- Elastic and factorised scattering







Assumptions

• $\theta_i \neq \theta_j, \, i \neq j$

 $\Rightarrow IR \text{ is finite} \\ \Rightarrow Every \text{ soliton interacts} \\ \text{with every other in the IR} \\ \end{cases}$



Assumptions

- $\theta_i \neq \theta_j, i \neq j$
- $\Rightarrow IR \text{ is finite} \\ \Rightarrow Every \text{ soliton interacts} \\ \text{with every other in the IR} \\ \end{cases}$

• Shift

$$K_{ij} \equiv K(v_i, \theta_i; v_j, \theta_j)$$
$$K_i = \sum_{j \neq i} K_{ij}$$



Assumptions

- $\theta_i \neq \theta_j, i \neq j$
- $\Rightarrow IR is finite$ $\Rightarrow Every soliton interacts$ with every other in the IR

• Shift

- $K_{ij} \equiv K(v_i, \theta_i; v_j, \theta_j)$ $K_i = \sum_{j \neq i} K_{ij}$
- Homogeneous gas in bulk of IR



(10)



• Geometric argument

 $\theta_i^{\text{eff}} = \theta_i - \arcsin\frac{K_i}{L_i}$

 L_i^-

 K_i

 (v_i, θ_i)

 L_i

 $(v_i^{\text{eff}}, \theta_i^{\text{eff}})$

 $(v_i, heta_i)$

• Geometric argument

$$\theta_i^{\text{eff}} = \theta_i - \arcsin\frac{K_i}{L_i}$$

• Introduce line density $\tilde{\rho}_i(v,\theta)$

$$K_i = \sum_j K_{ij} \approx L_i \int \tilde{\rho}_i(u, \alpha) K(v_i, \theta_i; u, \alpha) \mathrm{d}u \mathrm{d}\alpha$$

 L_i^-

 K_i

 (v_i, θ_i)

 L_i

 $(v_i^{\text{eff}}, \theta_i^{\text{eff}})$

 (v_i, θ_i)

• Geometric argument

$$\theta_i^{\text{eff}} = \theta_i - \arcsin\frac{K_i}{L_i}$$

• Introduce line density $\tilde{\rho}_i(v,\theta)$

$$K_i = \sum_j K_{ij} \approx L_i \int \tilde{\rho}_i(u, \alpha) K(v_i, \theta_i; u, \alpha) \mathrm{d}u \mathrm{d}\alpha$$

• Integral equation

 $\theta^{\text{eff}}(v,\theta) = \theta - \arcsin\left[\int \tilde{\rho}(u,\alpha;\theta^{\text{eff}}(v,\theta))K(v,\theta;u,\alpha)\mathrm{d}u\mathrm{d}\alpha\right]$

 L_i^-

 K_i

 (v_i, θ_i)

 L_i

 $(v_i^{\text{eff}}, \theta_i^{\text{eff}})$

 (v_i, θ_i)



$$\theta_i^{\text{eff}} = \theta_i - \arcsin\frac{K_i}{L_i}$$

• Introduce line density $\tilde{\rho}_i(v, \theta)$

$$K_i = \sum_j K_{ij} \approx L_i \int \tilde{\rho}_i(u, \alpha) K(v_i, \theta_i; u, \alpha) du d\alpha$$

• Integral equation

 $\theta^{\text{eff}}(v,\theta) = \theta - \arcsin\left[\int \tilde{\rho}(u,\alpha;\theta^{\text{eff}}(v,\theta))K(v,\theta;u,\alpha)\mathrm{d}u\mathrm{d}\alpha\right]$ $\tilde{\rho}(v,\theta;\theta_i^{\text{eff}}) = \tilde{\rho}_i(v,\theta)$

• Want to relate $\tilde{\rho}(v,\theta;\varphi)$ to $\tilde{\rho}(v,\theta;0)$

• Want to relate $\tilde{\rho}(v,\theta;\varphi)$ to $\tilde{\rho}(v,\theta;0)$

$$\tilde{\rho}(v,\theta;0) \sim \frac{1}{\Delta(v,\theta)}$$
 and $\tilde{\rho}(v,\theta;\varphi) \sim \frac{1}{\Delta'(v,\theta)}$

11

Typical distance between two solitons with parameters in the vicinity of (v, θ) as they intersect the horizontal

• Want to relate $\tilde{\rho}(v,\theta;\varphi)$ to $\tilde{\rho}(v,\theta;0)$

$$\tilde{\rho}(v,\theta;0) \sim \frac{1}{\Delta(v,\theta)}$$
 and $\tilde{\rho}(v,\theta;\varphi) \sim \frac{1}{\Delta'(v,\theta)}$

• Geometric argument



• Want to relate $\tilde{\rho}(v,\theta;\varphi)$ to $\tilde{\rho}(v,\theta;0)$

$$\Delta = \Delta' \sin \varphi \left(\cot \varphi - \cot \theta^{\text{eff}} \right)$$
$$\bigcup$$
$$\theta; \varphi) = \tilde{\rho}(v, \theta; 0) |\sin \varphi| \left| \cot \varphi - \cot \theta^{\text{eff}} \right|$$

• Want to relate $\tilde{\rho}(v,\theta;\varphi)$ to $\tilde{\rho}(v,\theta;0)$

$$\tilde{\rho}(v,\theta;0) \sim \frac{1}{\Delta(v,\theta)} \quad \text{and} \quad \tilde{\rho}(v,\theta;\varphi) \sim \frac{1}{\Delta'(v,\theta)}$$
• Geometric argument
$$\Delta = \Delta' \sin \varphi \left(\cot \varphi - \cot \theta^{\text{eff}}\right)$$

$$\overline{\phi}$$

$$\tilde{\rho}(v,\theta;\varphi) = \tilde{\rho}(v,\theta;0) |\sin \varphi| \left|\cot \varphi - \cot \theta^{\text{eff}}\right|$$

Effective orientation

 $\theta^{\text{eff}}(v,\theta) = \theta - \arcsin\left[\int \tilde{\rho}(u,\alpha;0)K(v,\theta;u,\alpha)\left|\sin\theta^{\text{eff}}(v,\theta)\right|\left|\cot\theta^{\text{eff}}(v,\theta) - \cot\alpha^{\text{eff}}(u,\alpha)\right|dud\alpha\right]$

 $\frac{1}{\Delta'(v,\theta)}$

- Similarities with refraction: there should be a way to relate $\theta^{\rm eff}$ and $v^{\rm eff}$



[12]

- Similarities with refraction: there should be a way to relate θ^{eff} and v^{eff}
- Geometric argument: assume that over dt IR does not change and over dx it can be considered flat



- Similarities with refraction: there should be a way to relate θ^{eff} and v^{eff}
- Geometric argument: assume that over dt IR does not change and over dx it can be considered flat



v^{eff}	=	$\sin \theta^{\mathrm{eff}}$
\overline{v}		$\sin \theta$

Snell's law!

- Similarities with refraction: there should be a way to relate θ^{eff} and v^{eff}
- Geometric argument: assume that over dt IR does not change and over dx it can be considered flat

$$\frac{v^{\text{eff}}}{v} = \frac{\sin \theta^{\text{eff}}}{\sin \theta}$$

Snell's law!

• Remark 1: θ w.r.t the horizontal

- Similarities with refraction: there should be a way to relate θ^{eff} and v^{eff}
- Geometric argument: assume that over dt IR does not change and over dx it can be considered flat

$$\frac{v^{\text{eff}}}{v} = \frac{\sin \theta^{\text{eff}}}{\sin \theta} \qquad \qquad \text{Snell's law!}$$

• Remark 1: θ w.r.t the horizontal

• Remark 2: for $\theta^{\text{eff}} \approx \theta$

$$v^{\text{eff}} \approx v \left\{ 1 - \cot \theta \left[\int \tilde{\rho}(u,\alpha;0) K(v,\theta;u,\alpha) \left| \sin \theta^{\text{eff}}(v,\theta) \right| \left| \cot \theta^{\text{eff}}(v,\theta) - \cot \alpha^{\text{eff}}(u,\alpha) \right| du d\alpha \right\} \right\}$$

• Gas of lines allows for analogy with (1+1)D systems



Space-time trajectories of quasi particles in a (1+1)d system

Snapshot of KP N-soliton solution in the (x,y) plane

• Gas of lines allows for analogy with (1+1)D systems



[Courtesy of G. Roberti]

- Snapshot of KP Nsoliton solution in the (x,y) plane
- At fixed time KP equation yields an (integrable) Boussinesq equation

 $6\partial_x(u\partial_x u) + \partial_{xxxx}u \pm 3\partial_{yy}u = 0$



• Gas of lines allows for analogy with (1+1)D systems



[Courtesy of G. Roberti] Snapshot of KP Nsoliton solution in the (x,y) plane

• At fixed time KP equation yields an (integrable) Boussinesq equation

 $6\partial_x(u\partial_x u) + \partial_{xxxx}u \pm 3\partial_{yy}u = 0$

• Proposition: study the gas of lines through GHD of its (1+1)D counterpart

• Gas of lines allows for analogy with (1+1)D systems



[Courtesy of G. Roberti] Snapshot of KP Nsoliton solution in the (x,y) plane

• At fixed time KP equation yields an (integrable) Boussinesq equation

 $6\partial_x(u\partial_x u) + \partial_{xxxx}u \pm 3\partial_{yy}u = 0$

• Proposition: study the gas of lines through GHD of its (1+1)D counterpart

 \Rightarrow Snapshot from (1+1)D space-time trajectories

• Gas of lines allows for analogy with (1+1)D systems



[Courtesy of G. Roberti] Snapshot of KP Nsoliton solution in

- the (x,y) plane
- At fixed time KP equation yields an (integrable) Boussinesq equation

 $6\partial_x(u\partial_x u) + \partial_{xxxx}u \pm 3\partial_{yy}u = 0$

• Proposition: study the gas of lines through GHD of its (1+1)D counterpart

 \Rightarrow Snapshot from (1+1)D space-time trajectories

 \Rightarrow Dynamics obtained by varying the impact parameters $x_{1D,i}^-$ via $v_{2D,i}$

Refraction revisited



(14)

Х






 θ

t



• Identify θ with v_{1D}

$$v_{1\mathrm{D}} \leftrightarrow \frac{1}{\tan \theta}$$
 and $v_{1\mathrm{D}}^{\mathrm{eff}} \leftrightarrow \frac{1}{\tan \theta^{\mathrm{eff}}}$

• Recall "Snell's law"

$$\frac{v_{2D}^{\text{eff}}}{v_{2D}} = \frac{\sin\theta^{\text{eff}}}{\sin\theta} = \sqrt{\frac{1 + \cot^2\theta}{1 + \cot^2\theta^{\text{eff}}}}$$

(14)



• Identify θ with v_{1D}

$$v_{1\mathrm{D}} \leftrightarrow \frac{1}{\tan \theta}$$
 and $v_{1\mathrm{D}}^{\mathrm{eff}} \leftrightarrow \frac{1}{\tan \theta^{\mathrm{eff}}}$

• Recall "Snell's law"

$$\frac{v_{\rm 2D}^{\rm eff}}{v_{\rm 2D}} = \frac{\sin\theta^{\rm eff}}{\sin\theta} = \sqrt{\frac{1+\cot^2\theta}{1+\cot^2\theta^{\rm eff}}}$$

 $\Rightarrow \quad (v_{2D}^{\text{eff}})^2 \left[1 + (v_{1D}^{\text{eff}})^2 \right] = v_{2D}^2 \left[1 + v_{1D}^2 \right]$





• Identify θ with v_{1D}

$$v_{1\mathrm{D}} \leftrightarrow \frac{1}{\tan \theta} \quad \text{and} \quad v_{1\mathrm{D}}^{\mathrm{eff}} \leftrightarrow \frac{1}{\tan \theta^{\mathrm{eff}}}$$

• Recall "Snell's law"

$$\frac{v_{2\mathrm{D}}^{\mathrm{eff}}}{v_{2\mathrm{D}}} = \frac{\sin\theta^{\mathrm{eff}}}{\sin\theta} \quad \Rightarrow \quad (v_{2\mathrm{D}}^{\mathrm{eff}})^2 \left[1 + (v_{1\mathrm{D}}^{\mathrm{eff}})^2\right] = v_{2\mathrm{D}}^2 \left[1 + v_{1\mathrm{D}}^2\right]$$

• Identify line density $\tilde{\rho}(v,\theta;0)$ with DOS $\rho_{1D}(p)$

$$\int \tilde{\rho}(v,\theta;0) dv d\theta = \int \rho_{1D}(p) dp$$
$$\Rightarrow \quad \rho_{1D}(p) = \frac{v_{1D}'(p)}{1 + v_{1D}^2(p)} \int \tilde{\rho}(v, \operatorname{acot} p; 0) dv$$



• Use GHD of a (1+1)D models with different types of particles so that

$$\rho_{1\mathrm{D}}(v,p) = \frac{v_{1\mathrm{D}}'(p)}{1 + v_{1\mathrm{D}}^2(p)} \tilde{\rho}(v, \operatorname{acot} \, p; 0)$$

15

• Use GHD of a (1+1)D models with different types of particles so that

$$\rho_{1\mathrm{D}}(v,p) = \frac{v_{1\mathrm{D}}'(p)}{1 + v_{1\mathrm{D}}^2(p)} \tilde{\rho}(v, \operatorname{acot} \, p; 0)$$

• Relate phase shifts from (2+1)D model and that of its (1+1)D counterpart

• Use GHD of a (1+1)D models with different types of particles so that

$$\rho_{1\mathrm{D}}(v,p) = \frac{v_{1\mathrm{D}}'(p)}{1 + v_{1\mathrm{D}}^2(p)} \tilde{\rho}(v, \operatorname{acot} \, p; 0)$$

- Relate phase shifts from (2+1)D model and that of its (1+1)D counterpart
- Partition function of the gas of lines from that of its (1+1)D counterpart

• Use GHD of a (1+1)D models with different types of particles so that

$$\rho_{1\mathrm{D}}(v,p) = \frac{v_{1\mathrm{D}}'(p)}{1 + v_{1\mathrm{D}}^2(p)} \tilde{\rho}(v, \operatorname{acot} \, p; 0)$$

- Relate phase shifts from (2+1)D model and that of its (1+1)D counterpart
- Partition function of the gas of lines from that of its (1+1)D counterpart

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}v_{1\mathrm{D}}(p_i)}{2\pi} \mathrm{d}v_i \mathrm{d}x_{1\mathrm{D},i}^- \exp\left[-\sum_{i=1}^N w(v_i,\theta_i)\right] \chi\left(u_{1\mathrm{D},N}(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

• Use GHD of a (1+1)D models with different types of particles so that

$$\rho_{1\mathrm{D}}(v,p) = \frac{v_{1\mathrm{D}}'(p)}{1 + v_{1\mathrm{D}}^2(p)} \tilde{\rho}(v, \operatorname{acot} \, p; 0)$$

- Relate phase shifts from (2+1)D model and that of its (1+1)D counterpart
- Partition function of the gas of lines from that of its (1+1)D counterpart

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}v_{1\mathrm{D}}(p_i)}{2\pi} \mathrm{d}v_i \mathrm{d}x_{1\mathrm{D},i}^- \exp\left[-\sum_{i=1}^N w(v_i,\theta_i)\right] \chi\left(u_{1\mathrm{D},N}(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

15

• Correlations, entropy, σ_{2D} ?

• Use GHD of a (1+1)D models with different types of particles so that

$$\rho_{1\mathrm{D}}(v,p) = \frac{v_{1\mathrm{D}}'(p)}{1 + v_{1\mathrm{D}}^2(p)} \tilde{\rho}(v, \operatorname{acot} \, p; 0)$$

- Relate phase shifts from (2+1)D model and that of its (1+1)D counterpart
- Partition function of the gas of lines from that of its (1+1)D counterpart

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{1}{N!} \int_{\Gamma^N \times \mathbb{R}^N} \prod_{i=1}^N \frac{\mathrm{d}v_{1\mathrm{D}}(p_i)}{2\pi} \mathrm{d}v_i \mathrm{d}x_{1\mathrm{D},i}^- \exp\left[-\sum_{i=1}^N w(v_i,\theta_i)\right] \chi\left(u_{1\mathrm{D},N}(x,t=0) < \epsilon_x, x \notin [0,L]\right)$$

- Correlations, entropy, σ_{2D} ?
- For any (1+1)D integrable model there should be an associated gas of lines. How much is this relevant ?